UNIVERSITÄT LEIPZIG

Climate Dynamics (Summer Semester 2018) J. Mülmenstädt

Today's Lecture (Lecture 10): Internal variability

Reference

IPCC AR5, Ch. 14 Holton, Ch. 11 COMET Introduction to tropical meteorology, linked from course webpage

3 – Internal variability

1. Introduction

2. The climate system

3. Internal variability

3.1 Departures from temporal average
3.2 Modes of internal variability
3.3 El Niño-Southern Oscillation
3.4 Modeling ENSO
3.5 ENSO teleconnections

4. Forcing and feedbacks

5. Anthropogenic climate change

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3.1 – Departures from temporal average



Classification of fluctuations about the climatological mean:

- Time scale: for example, diurnal, annual cycles in solar forcing
- Periodic or trend?
- Natural or anthropogenic?
- Forcing or internal variability?

Forcing or internal variability?

Forcing

- Change in net energy exchange between the climate system and the environment
- Examples:
 - Solar intensity cycles
 - Orbital parameter cycles
 - Volcanic eruptions
 - Anthropogenic emissions of greenhouse gases, aerosols

Internal variability

- Many aspects of the climate system
 - are described by nonlinear differential equations
 - couple systems with different time scales
 - have positive feedback
- This leads to transitions between states (example: the attractors in the Lorenz system) on various time scales

- Examples (with time scales):
- Days-weeks Midlatitude storm systems
 - Months Madden-Julian oscillation (MJO)
- Interannual El Niño-southern oscillation (ENSO)
 - Decadal Pacific decadal oscillation (PDO)
- Multidecadal Atlantic multidecadal oscillation (AMO)

3.2 – Modes of internal variability

What is a mode?

A *mode* describes the space-time structure of the variability. Often it is defined as a product of a characteristic spatial *pattern* (of one or more variables) and a time-varying *index*. In next week's homework, we will have a hands-on exercise on determining a mode.



Modes of internal variability



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3.3 - El Niño-Southern Oscillation

- > ENSO describes cyclical changes in the state of the tropical Pacific Ocean and the overlying atmosphere
- > Through teleconnections, ENSO influences not only the tropical Pacific but most other regions, including the Arctic
- > Strong influence on regional climates (temperature and precipitation) and on global-mean temperature
- > On interannual time scales, ENSO is the most important mode of variability globally
- Strong El Niños are visible in the global-mean temperature time series (e.g., 1998)



ENSO and global mean surface temperature



Figure: NASA GISS

El Niño-Southern Oscillation

Southern oscillation – an atmospheric phenomenon









But the underlying process is a combined ocean-atmosphere phenomenon (ENSO)

Mean state of the equatorial atmosphere – Walker circulation



Figure: Holton

Oceanic and atmospheric circulation are coupled through convection

December - February Normal Conditions





Mean state of the equatorial ocean - SST, thermocline depth, sea surface elevation



El Niño-La Niña cycle



Walker circulation: normal (top) and El Niño (bottom)





Figure: Holton

3.4 – Modeling ENSO

- A satisfactory model of ENSO must explain
 - 1. Formation of the El Niño pattern
 - 2. Collapse of El Niño and transition to La Niña
 - 3. Collapse of La Niña pattern and return to neutral conditions

There are many models that accomplish this (more or less).

> The irregular temporal cycle (which is in part due to interactions with other modes) is hard to forecast

- > For climate models, the important features are
 - Realistic time spectrum
 - Realistic teleconnections (export of energy from tropical ocean)

The purpose of conceptual models is to simplify the ENSO problem to just those characteristics that explain the observed behavior. Two examples of conceptual models for ENSO are:

Delayed oscillator

explains why La Niña follows El Niño

Recharge-discharge of ocean heat

explains why El Niño causes a spike in global mean temperature (by releasing heat stored in the ocean)

Some tropical dynamics – tropical waves

Rossby-gravity wave velocity anomaly field



Kelvin wave velocity anomaly field



These wave disturbances (horizontal velocity perturbations sketched in the figures) are *equatorially trapped* (group velocity is zonal along the equator)

Figure: Holton, Ch. 11

Equatorial wave theory

Near equator, to good approximation (equatorial beta plane):

$$F = \beta y$$
 (3.1)

For an incompressible fluid of mean depth h_e , we can write the following linearized set of equations for perturbation u, v and Φ :

$$\frac{\partial u'}{\partial t} - \beta y v' = -\frac{\partial \Phi'}{\partial x} \qquad \text{equation of motion} \qquad (3.2)$$

$$\frac{\partial v'}{\partial t} + \beta y u' = -\frac{\partial \Phi'}{\partial y} \qquad \text{note: not geostrophic at equator} \qquad (3.3)$$

$$\frac{\partial \Phi'}{\partial t} + g w' = -g h_e \left(\frac{\partial u'}{\partial x} + \frac{\partial v'}{\partial y}\right) \qquad \text{hydrostatic equilibrium + continuity} \qquad (3.4)$$

These linearized equations assume that $\overline{u} = 0$, $u' \partial u' / \partial x \ll \partial u' / \partial t$ etc. (terms second-order and higher in the perturbation \ll first-order). They apply to ocean and atmosphere.

Equatorial wave theory

Ansatz for zonally propagating waves:

$$\begin{pmatrix} u'\\ v'\\ \Phi' \end{pmatrix} = \begin{pmatrix} \hat{\nu}(y)\\ \hat{\nu}(y)\\ \hat{\Phi}(y) \end{pmatrix} \exp(i(kx - \nu t))$$
(3.5)

Insert into (3.2)- (3.4):

$$-i\nu\hat{\boldsymbol{\upsilon}} - \beta \mathbf{y}\hat{\boldsymbol{\upsilon}} = -ik\hat{\boldsymbol{\Phi}}$$
(3.6)

$$-i\nu\hat{\mathbf{v}} + \beta \mathbf{y}\hat{\mathbf{v}} = -\frac{\partial\hat{\mathbf{\Phi}}}{\partial \mathbf{y}}$$
(3.7)

$$-i\nu\hat{\Phi} + gh_e\left(ik\hat{\upsilon} + \frac{\partial\hat{\upsilon}}{\partial y}\right) = 0$$
(3.8)

which yields two first-order differential equations in two variables (after elimination of \hat{v}):

$$\left(\beta^2 y^2 - \nu^2\right)\hat{\mathbf{v}} = ik\beta y\hat{\Phi} + i\nu\frac{\partial\hat{\Phi}}{\partial y}$$
(3.9)

$$\left(\nu^{2} - gh_{e}k^{2}\right)\hat{\Phi} + i\nu gh_{e}\left(\frac{\partial\hat{v}}{\partial y} - \frac{k}{\nu}\beta y\hat{v}\right) = 0$$
(3.10)

or

$$\frac{\partial^2 \hat{v}}{\partial y^2} + \left(\frac{\nu^2}{gh_e} - k^2 - \frac{k}{\nu}\beta - \frac{\beta^2 y^2}{gh_e}\right)\hat{v} = 0$$
(3.11)

(3.11) has solutions of the form

$$v(\xi) = v_0 H_n(\xi) \exp\left(-\frac{\xi^2}{\nu}\right), \quad \text{where} \quad \xi = \left(\frac{\beta}{\sqrt{gh_e}}\right)^{\frac{1}{2}} y, \quad n = 0, 1, 2, \dots$$
(3.12)

and H_n are the hermite polynomials,

$$H_0(\xi) = 1, \quad H_1(\xi) = 2\xi, \quad H_2(\xi) = 4\xi^2 - 2$$
 (3.13)

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Equatorial wave theory

Rossby-gravity waves

Consider the n = 0 solution. Substituting this \hat{v} into (3.11) leads to a dispersion relation:

$$\nu = k\sqrt{gh_e} \left(\frac{1}{2} \pm \frac{1}{2}\sqrt{1 + \frac{4\beta}{k^2\sqrt{gh_e}}}\right)$$
(3.14)

This type of wave is called *Rossby-gravity wave*. The two solutions (3.14) have opposite signs of group velocity $(c_g = \partial \nu / \partial k)$ and phase velocity $(c = \nu / k)$, representing a westward-traveling and eastward-traveling wave.

Kelvin waves

Note that $\hat{v} = 0$ is a (trivial) solution of (3.11). This solution is called a *Kelvin wave*. In this case the system of equations (3.6)–(3.8) simplifies to

$$\beta \mathbf{y} \hat{\mathbf{v}} = -\frac{\nu}{k} \frac{\partial \hat{\mathbf{v}}}{\partial \mathbf{y}} \tag{3.15}$$

which has the solution

$$\hat{u} = u_0 \exp\left(-\frac{\beta y^2}{2c}\right) \tag{3.16}$$

From (3.16) it is evident that only Kelvin waves with $c = \nu/k > 0$, i.e., eastward-traveling Kelvin waves, are equatorially trapped.

Dispersion relations



Simple Model of Wind Induced Perturbation of the Tropical Pacific Ocean

Delayed oscillator model: ENSO disturbance modeled as Kelvin and Rossby-gravity waves

4. 25 days 125 days Intribute Langitud 50 days 175 days Lansitud Lansitude 3 5 75 days 225 days Longitude 100 days 6 275 days Lonsitud

Time Evolution for the Idealized Experimental Kelvin and Rossby Waves Across the Pacific

- Downwelling Kelvin wave (KW) forms in response to westerly wind anomaly; KW propagates eastward with 2–3 m s⁻¹ speed (2–3 months to cross Pacific)
- 2. The downwelling KW deepens the thermocline and warms the eastern Pacific
- 3. Upwelling Rossby wave (RW) propagates westward (1 m s⁻¹), leading to a shallower thermocline in the western Pacific
- 4. Reflection of the upwelling RW leads to upwelling eastward-propagating KW + RW
- 5. After \sim 8 months, the upwelling KW has crossd the Pacific and reversed the thermocline deepening in the eastern Pacific
- KW at eastern boundary is partly reflected as RW, partly as KW. Recall that only eastward-propagating KW are equatorially trapped; the reflected KW propagates poleward (as coastal KW)

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3.5 - ENSO teleconnections: they are nearly global



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Tropical teleconnections include cyclone frequency



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